



Physics of Extreme Massive Stars

Marie-Curie-RISE project funded by the European Union



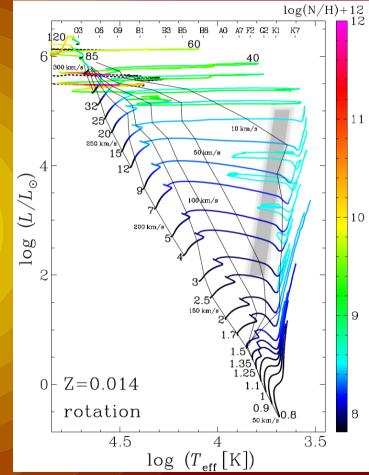
The mass-loss activity of yellow hypergiants

Michalis Kourniotis

Astronomical Institute Czech Academy of Sciences

Summer school on Stellar Winds and Outflows Harrachov – 13/09/2023

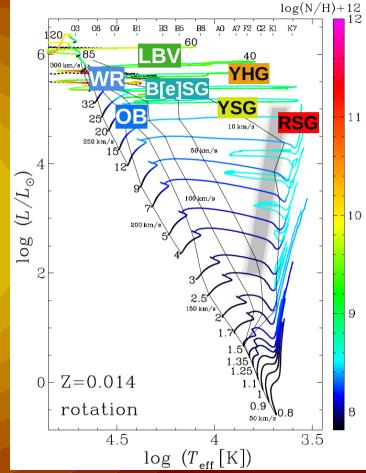
Massive star evolution



- Massive stars (> 8 M_o) display short but impacting life
- Different channels give rise to the various stellar types
- The mass-loss mechanisms vary throughout evolution; their physics are associated to the stellar parameters
- Post-RSG evolution remains a missing piece of the puzzle

Ekström et al. 2012

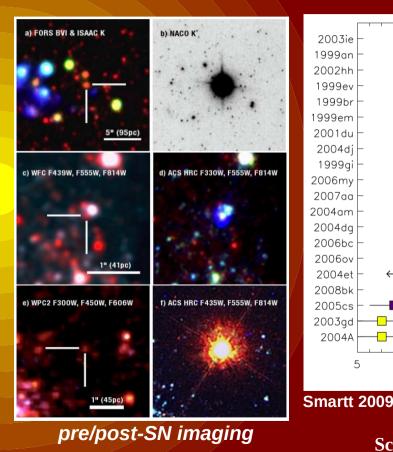
Massive star evolution



Ekström et al. 2012

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- Post-RSG evolution remains a missing piece of the puzzle

The RSG problem or... how do RSGs with M > 17 M_o die ?

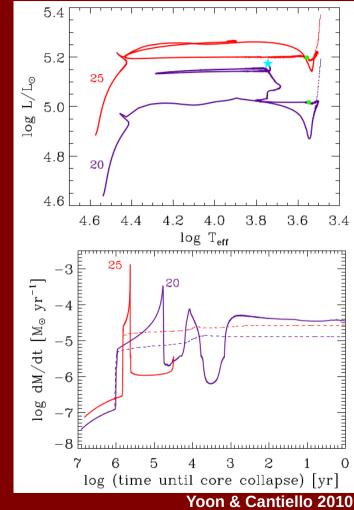


School on Stellar Winds and Outflows

Initial mass / M_o

 \leftarrow

High mass-losses during RSG phase



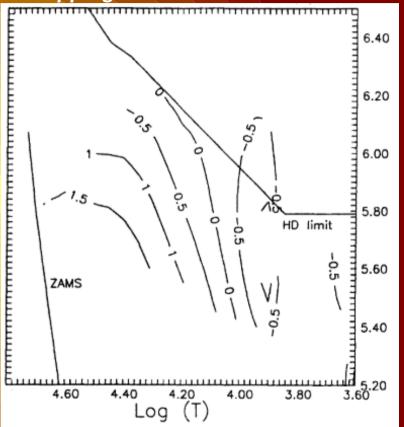
- 10

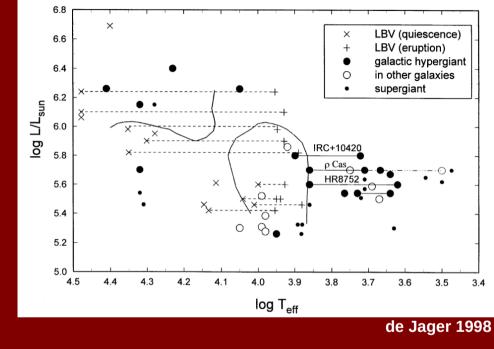
- 9

- 8 - 7

Atmospheric instability

Mapping the effective acceleration





Regions of atmospheric instability

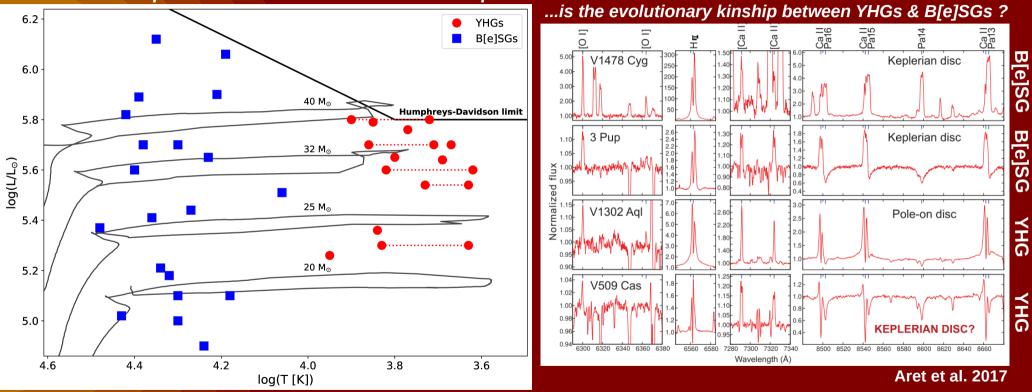
 \checkmark The "Yellow Void" (YV) is modeled as area of very low $g_{_{eff}}$ values

- ✓ It is situated within T_{eff} ~ 6000–10000 K with log(L/L_o) > 5.5
- YHGs are believed to exhibit outbursts when approaching YV

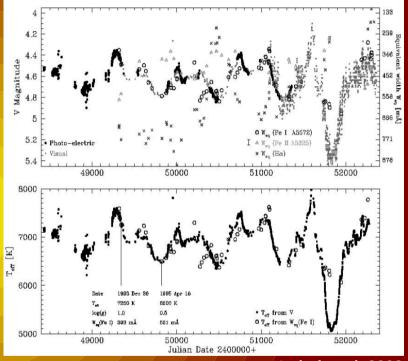
Nieuwenhuijzen et al. 1995

The post-RSGs: YHGs and the B[e]SGs

A census of post-RSGs classes in the Local Group

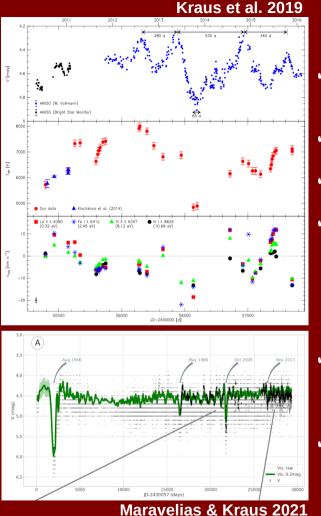


The prototype ρ Cas



Lobel et al. 2003

- Ejection of a 3x10⁻² M_o shell
- Formation of TiO lines during the outburst
- Extended, velocity-stratified atmosphere



School on Stellar Winds and Outflows

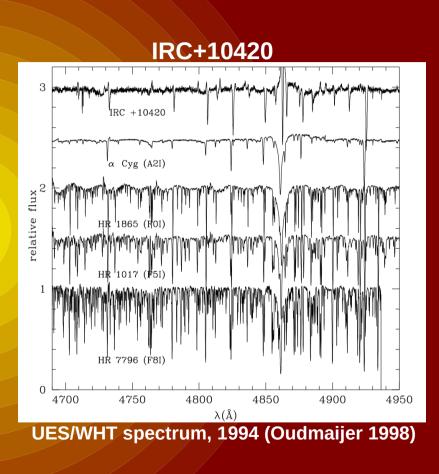
Temperature decrease of ~3000 K in 2013

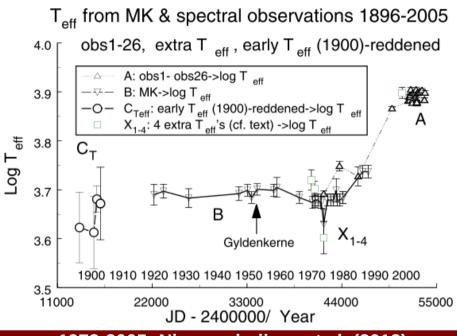
- Asymmetric lines
- Circumstellar emission
- Evidence of gas remnant from RSG phase

- The star underwent its most dramatic recorded outburst in 1946
- Decreasing duration and increasing frequency of the events

Other YHG variables

HR8752

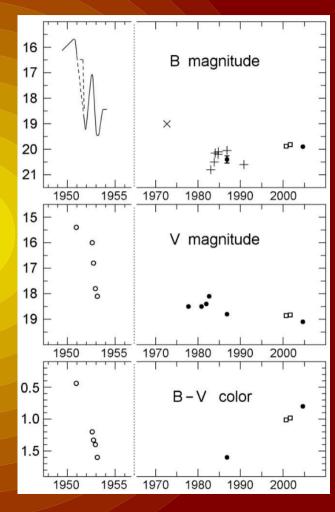


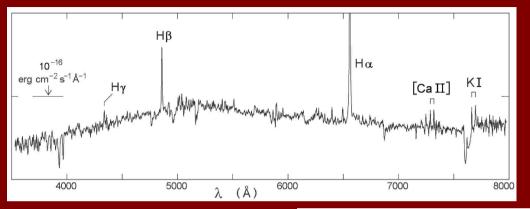


1973-2005; Nieuwenhuijzen et al. (2012)

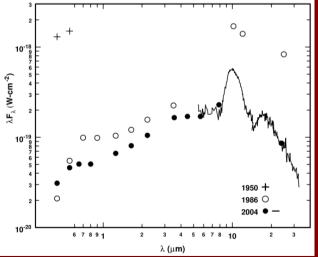
- IRC+10420: Change in spectral type from F8 to A2, blueshifted emission, and high extinction
- HR8752: The dispersion of past ejected shells in 1973 revealed the compact hot photosphere

Beyond the Galaxy: Var A



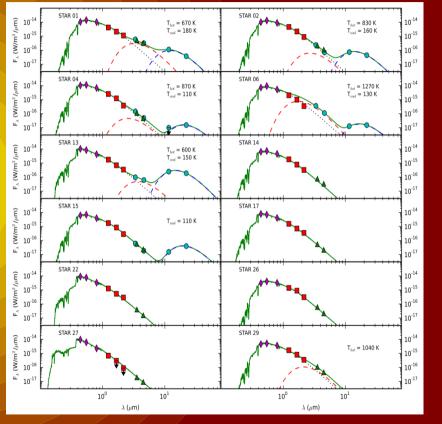


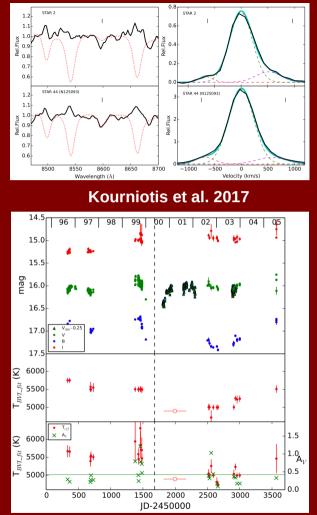
- Change in the spectral type from F to M
- Variability with timescale
 of > 50 years
- Heavy dust obscuration
 - Humphreys et al. 2006



Searching for YHGs

Exploring the surroundings of YHG candidates



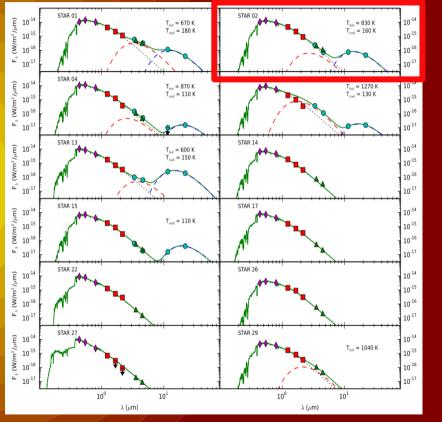


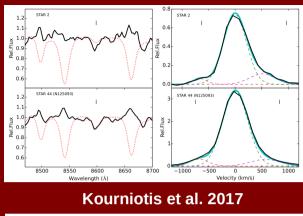


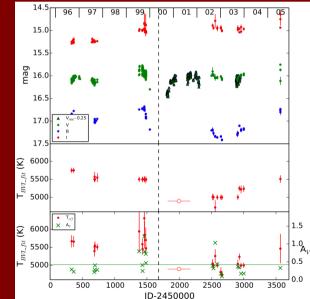
✓ Dimming event
 ✓ Drop by ∆T_{eff} > 500 K

Searching for YHGs

Exploring the surroundings of YHG candidates



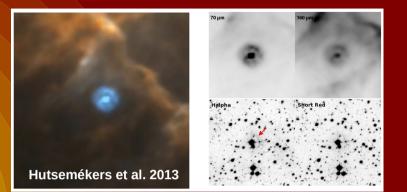




ν Broad Hα wings *ν* Call emission

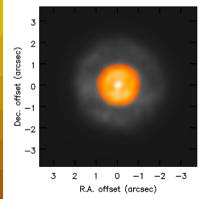


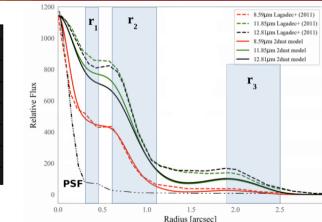
Nebulosity and shells of YHGs: Hen 3-1379

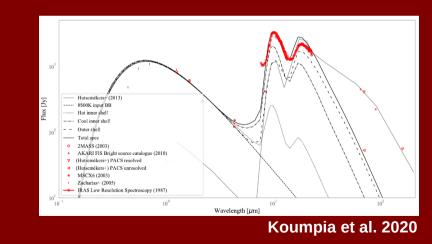


The Fried Egg nebula

Fried Egg Nebula	$\frac{M_{\rm gas}}{(10^{-3}M_\odot)}$	<i>T</i> _d (K)	r _d ('')	t _{kin} (yr)	\dot{M} $(M_{\odot}{ m yr}^{-1})$	
Hot inner shell	0.021	620–480	0.3-0.45	30.8	6×10^{-7}	
Intermediate shell	0.90	460-320	0.6 - 1.1	102.7	9×10^{-4}	
Outer shell	5.6	240-200	1.9-2.5	123.2	5×10^{-5}	
	$\log(L_{\star}/L_{\odot})$	T_{\star}	d			
IRAS 17163		(K)	(kpc)			
	5.7	8500	1.2			



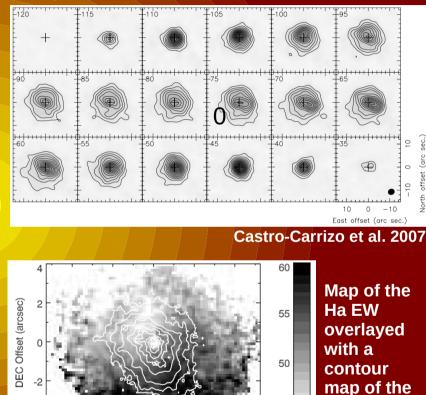


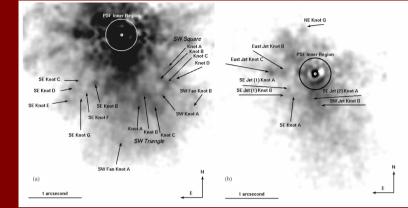


Koumpia et al. 2020

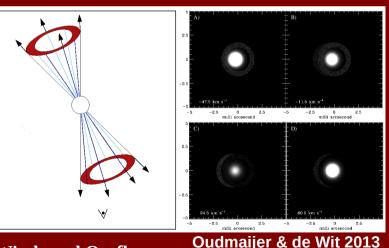
Nebulosity and shells of YHGs: IRC+10420

Maps of the 12CO J = 1–0 emission from IRAM





Tiffany et al. 2010



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B-band HST

image

45

HST/WFPC2 imaging

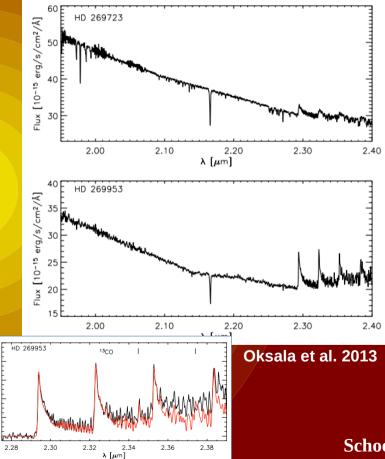
Geometry model based on AMBER observations

Davies et al. 2007

RA Offset (arcsec)

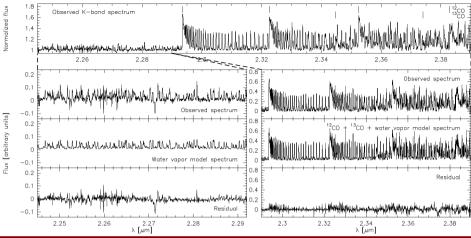
Ejecta in the K-band

Processed material around YHGs



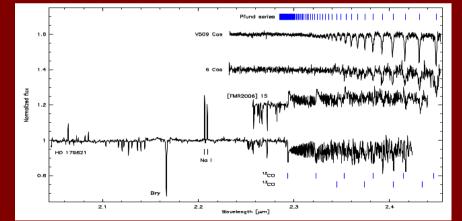
A

First detection of water vapor around HD269953



Kraus et al. 2022

Kraus et al. 2023

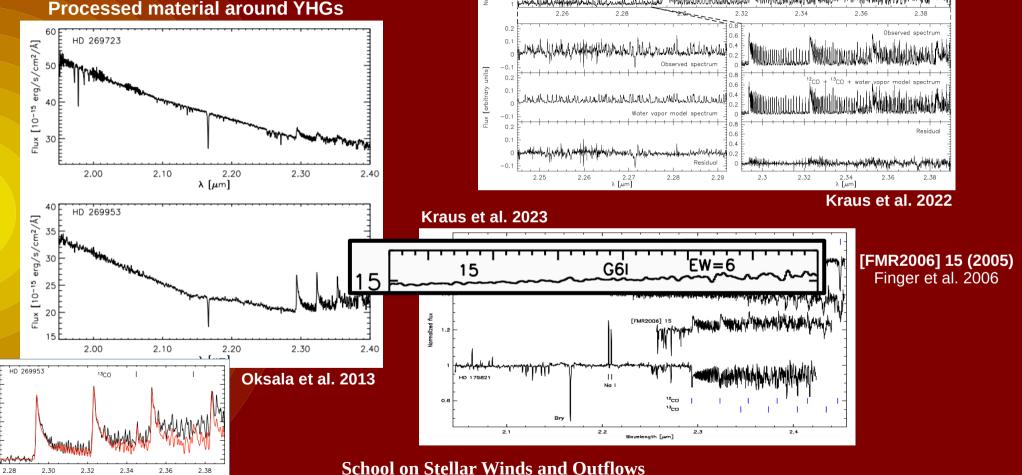


Ejecta in the K-band

Processed material around YHGs

A

λ [μm]



1.6 1,4 1.2

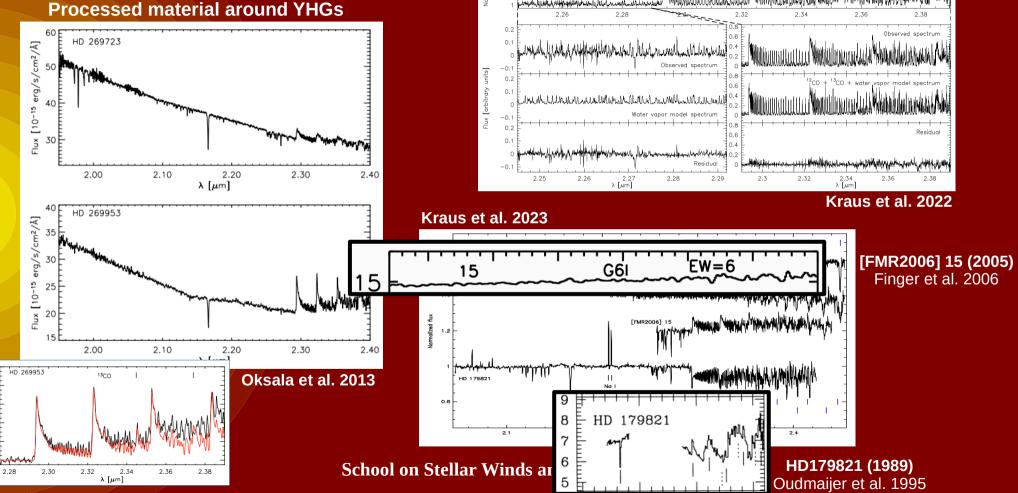
Observed K-band spectrum

First detection of water vapor around HD269953

Ejecta in the K-band

Processed material around YHGs

Ā



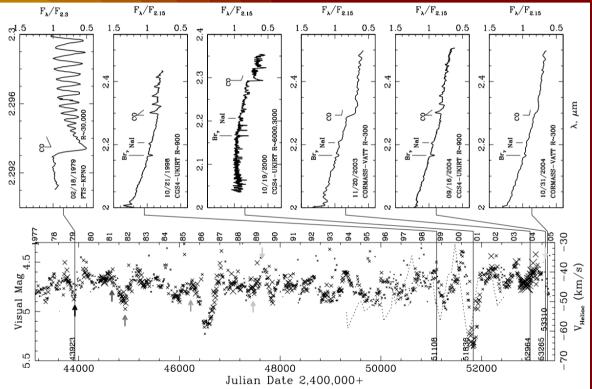
1,4 1.2

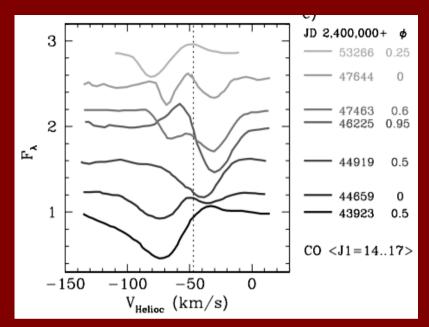
Observed K-band spectrum

First detection of water vapor around HD269953

Circumstellar variability

K-band spectroscopy pairing the photometric lightcurve

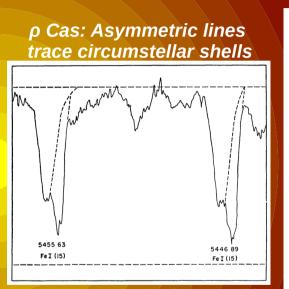




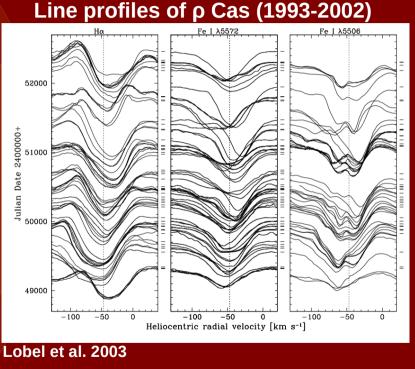
- ✓ A variable K-band appearance
- Correlates with stellar activity
- CO is found to be in emission at high mass-loss phases
- Line profile of a pulsating photosphere

Gorlova et al. 2006

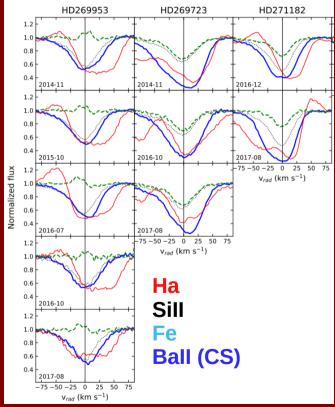
Line dynamics in the optical



Sargent 1961



Lines of YHGs in the LMC



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Kourniotis et al. 2022

Line dynamics in the optical

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51000+ 51000

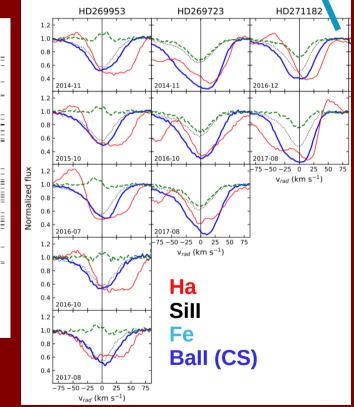
49000

Date

Julian J



Lines of YHGs in the LMC



ρ Cas: Asymmetric lines trace circumstellar shells

Sargent 1961

Lobel et al. 2003

-100

-50

0

School on Stellar Winds and Outflows

-50

Heliocentric radial velocity [km s⁻¹]

0

-100

-100

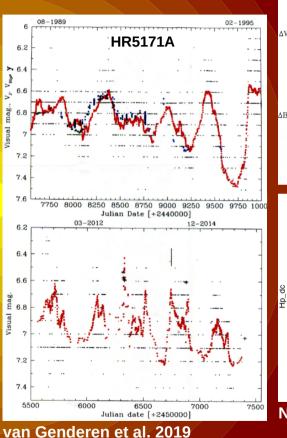
-50

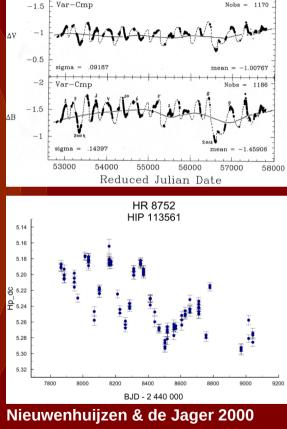
0

Line profiles of p Cas (1993-2002)

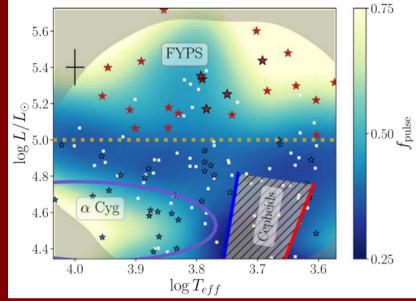
Kourniotis et al. 2022

Pulsational variability





rho Cas (2003-2017)

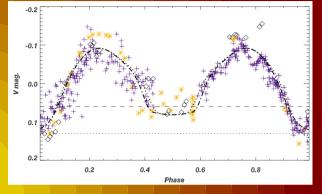


Dorn-Wallenstein et al. 2022

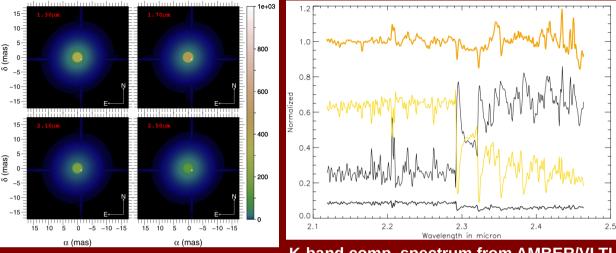
- Most, if not all, YHGs exhibit semi-regular pulsations with poorly-understood nature
- Pulsational properties change based on the location/direction of the YHG on the HRD
- How does the pulsating activity change in YHGs compared to YSGs ?

Binarity: HR5171A

Phased LCs (HIPPARCHOS/ASAS)



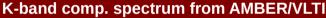
2



Models fitting AMBER/VLTI interf/ry

1700

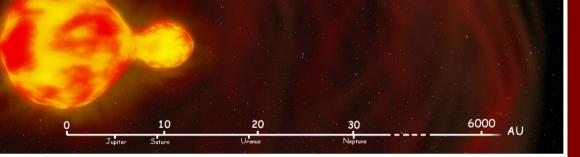
mas





- Extended nebulosity / flux asymmetry
- Eclipsing LC reveals system (P_{orb}~3.5yr)
- Possibly in Wind Roche-Lobe Overflow

Chesneau et al. 2014



Roche-lobe model of HR 5171A seen at phase 0.25

The astroquery package

hands on

- *astroquery* is a package for querying astronomical web forms and databases.
 Contains modules for querying web services for MAST, Simbad, Gaia, ESO, NASA ADS, Vizier etc, bringing the web interface to the scripting level.
- <u>HANDS-ON</u>: learning to query the Vizier service getting access to the published photometric catalogs, building the SED of a star.
 - Query the 2MASS catalog around the coordinates of YHG V382 Car
 - Converting magnitudes into fluxes using passband properties available from http://svo2.cab.inta-csic.es/theory/fps3/
 - Student task: <u>complement with Gaia 2 and WISE observations to assemble the</u> <u>optical/IR SED</u>
 - Student task: build SEDs for YHGs V509 Cas, IRC+10420, and ρ Cas

hands on

Radma IR byty X Y 2MASS All-Sky Catalog of Point Sources (Cutri+ 2003) II/246 Post annotation											
1.II/246/out	Filter ID	λ _{ref}	λ _{mean}	λ _{eff}	λ _{min}	λ _{max}	W _{eff}	ZΡ _v			
	2MASS/2MASS.J	12350.00	12350.00	12350.00	10806.47	14067.97	1624.32	1594.00			
	2MASS/2MASS.H	16620.00	16620.00	16620.00 16620.00		18231.02	2509.40	1024.00			
	2MASS/2MASS.Ks	21590.00	21590.00	21590.00	19543.69	23552.40	2618.87	666.80			
Radmin IR Optuvity WISE All-Sky Data Release (Cutri + 2012) II/311 Post annotation											
1.II/311/wise	Filter ID	λ_{ref} λ_n	nean)	eff	λ _{min}	λ _{max}	W _{eff}	ZΡ _ν			
	WISE/WISE.W1	33526.00	33526.00	33526.00	27540.97	38723.88	6626.42	309.54			
	WISE/WISE.W2	46028.00	46028.00	46028.00	39633.26	53413.60	10422.66	171.79			
	WISE/WISE.W3	115608.00	115608.00	115608.00	74430.44	172613.43	55055.23	31.67			
	WISE/WISE.W4	220883.00	220883.00	220883.00	195200.83	279107.24	41016.80	8.36			
Radmm R Opt uv X Y Gaia DR2 (Gaia Collaboration, 2018) acknowle I/345 1 annotation(s) (on 1 specific record(s))											
1.I/345/gaia2	Filter ID	λ _{ref}	λ _{mean}	λ _{eff}	λ _{min}	λ _{max}	W _{eff}	ZP _ν			
11/040/90102	GAIA/GAIA2.Gbp	5050.00	5279.89	5050.00	3280.45	6719.03	2347.38	3534.74			
GAIA/GAIA2.G 6230.00 6742.51 6230.00 3306.60 10450.65 41											
GAIA/GAIA2.Grp 7730.00 7883.71 7730.00 6254.97 10605.79 2756.78 26											
		School o	on Stellar Wi	nds and Out	flows						

The AAVSO database

hands on

The American Association of Variable Star Observers (1911) is an international nonprofit organization of observers, which enables anyone to participate in the scientific discovery in the field of variable star astronomy.

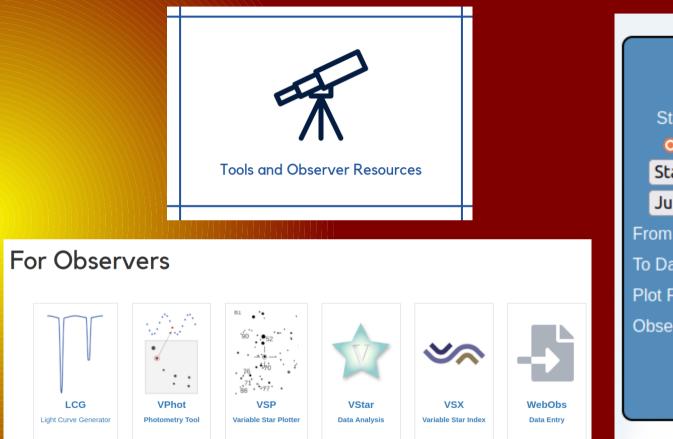
 It is based in Cambridge, Massachusetts, with active participants from more than 100 countries, hosting more than 34 million observations of variable stars.

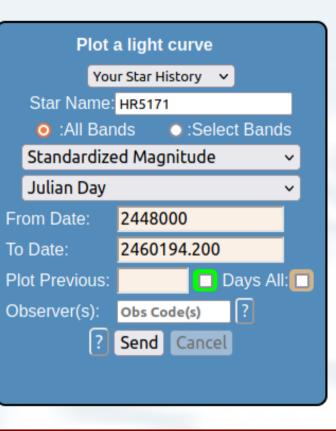
https://www.aavso.org/

<u>HANDS-ON</u>: learning to query, visualize and analyze a light curve from the AAVSO database

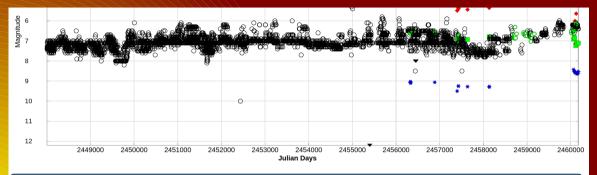
- Online lightcurve generator Accessing the data
- VStar tool for analysis of time-series
- Student task: periodic analysis of Eta Aql, ASAS 182611+1212.6, T UMi

hands on





hands on



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Alphabetical v

	AAP (5) : Vis	Abbott, Patrick	CA			AAX (256) : Vis
	ACN (1) : Vis	Adib, Carlos	BR	REA		AJC (8) : Vis
	ANAB (1) : Vis	Arias, Nicolás	AR			ARL (1) : Vis
	AWY (1) : Vis	Araujo, Wesley	BR			BWZ (216) : Vis
	BYU (22) : Vis	Brandie, J.	NZ			CFL (5) : Vis
	COM (15) : Vis	Cooper, Tim	ZA			CR (2) : Vis
	DCMA (1) : Vis	da Silva, Cledison	BR	AAVSO		DSI (46) : B,H,I,J,R,V
	DVC (2) : Vis	Devilliers, Chris	ZA			DXX (18) : Vis
	FMX (45) : Vis	Farrell, Fraser	AU			GAJ (13) : Vis
	GED (9) : Vis	Gonçalves, Eduardo Henrique Cordeiro	BR	AAVSO		GFE (1) : Vis
	GPX (53) : Vis	Goltz, William	AU			HEN (73) : Vis

Almeida, Joao BR Alencar Caldas, Romual BR NZ Blown, Eric Mean Points & Curve Curve: Smooth v : Mean Error Bars Visual (5679) V Bin Size(Julian Davs); 4.5 5.0 5.5 6.0 Magnitude 2.0 7.5 8.0 8.5 9.0

School on Stellar Winds and Outflows

BR REA

Amorim, Alexandre

hands on



Home

Download Data

Please complete this form to download variable star data from the AAVSO International For very large datasets or further assistance, please contact AAVSO HQ (aavso@aav WHAT IS THE AUID, NAME, OR DESIGNATION OF THE OBJECT? *

V766 CEN

WHAT DATE RANGE WOULD YOU LIKE TO USE? Enter Julian date or Gregorian Date (MM/DD/YYYY) START DATE *

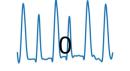
2448000

END DATE *

2460155.5

HOW SHALL WE CONTACT YOU? FIRST NAME

LAST NAME



Submit and Access Data

Variable Star Plotter

VSP Help Guide Request a Sequence Report chart errors Standard

PLOT A QUICK CHART

WHAT IS THE NAME, DESIGNATION OR AUID OF THE OBJECT?

V766 CEN

Required if no coordinates are provided below

RIGHT ASCENSION

Allowed Formats: HH:MM:SS, HH MM SS, DDD.XXXX. Required if no name is given above

CHOOSE A PREDEFINED CHART SCALE

Select one...

A is larger, slower; G is smaller, faster

CHOOSE A CHART ORIENTATION

Visual O Reversed O CCD

PLOT A FINDER CHART OR A TABLE OF FIELD PHOTOMETRY?*

O Chart O Photometry

CHART ID

A Chart ID will allow you to reproduce prior charts. Overrides all other fields in this form.

ADVANCED OPTIONS

FIELD OF VIEW

300

In Arcminutes. Must be between 0' and 1200

MAGNITUDE LIMIT

10.0



Variable Star Plotter

Plot Another Chart Photometry Table for This Chart

